# Palomar Observatory's

# **48-inch Samuel Oschin Telescope**

Owned & Operated by California Institute of Technology

By: Sam Pitts

Palomar Observatory Docent Non-Profit Educational Use

samsastro.com April 7, 2025

# 1928

Hale applies to International Education Board (IEB) to fund the 200-inch telescope with a grant from the Rockefeller Foundation

# **IEB** approves 6 Million Dollar Grant to Caltech

- A 200-inch reflecting telescope
- **B** a site, including land and land improvements
- **c** an Observatory & other necessary buildings
- **D** auxiliary apparatus
- E other expenses in connection with making the Observatory available for use

# The Worlds Largest Telescope needs Help Smaller Widefield Telescopes Capable of wide-angle images of the Night Sky

No detailed star charts or catalogues Need for a very Wide-Field Imaging Telescope It needs to go deep 16-20<sup>th</sup> Magnitude **Enter** 

# **1930 Barnard Schmidt**

### **The Schmidt Camera**

## Bernard Schmidt invents Schmidt Telescope1930



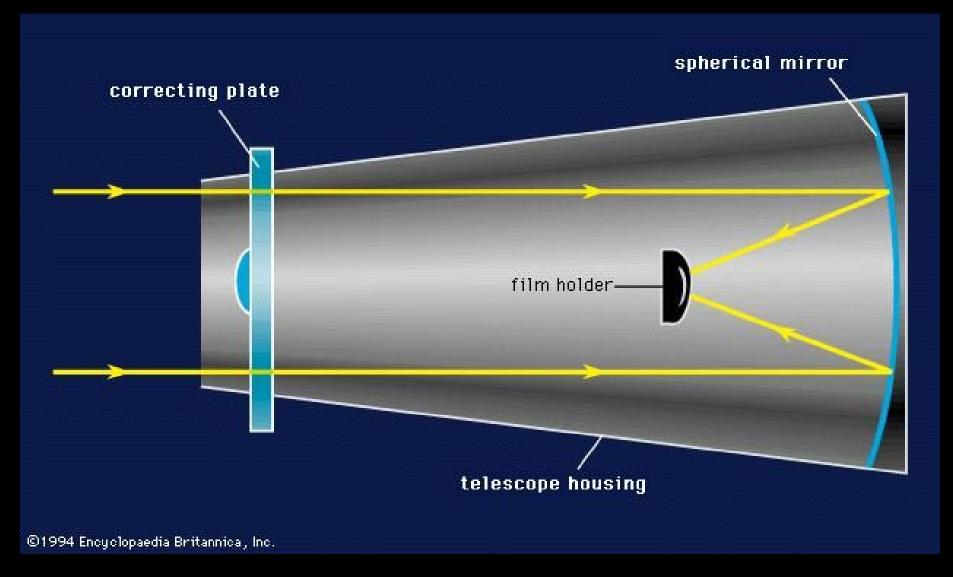
### March 30, 1879, – December 1, 1935

# The Schmidt Camera 1930

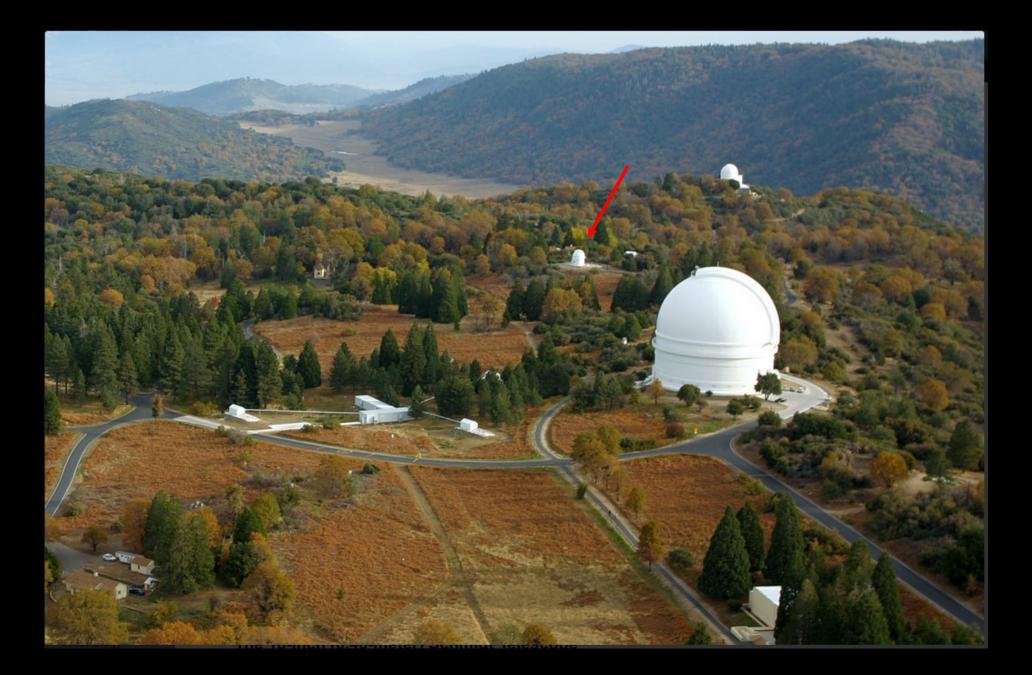
 $1^{st}$  Schmidt Camera  $14 \frac{1}{2}$  " (360mm) f/ 1.75 Main mirror of 44 cm (17.3 inch) Corrector plate of 36 cm (14 inch). The focal ratio was f=1.75, the field of view 7.5°

Corrector plate: Warping a parallel glass plate under partial vacuum into a slight sagging curve and then polishing the upper curve flat

# **The Schmidt Camera**



Large swathes of sky with short exposures in minutes versus hours with a reflector.

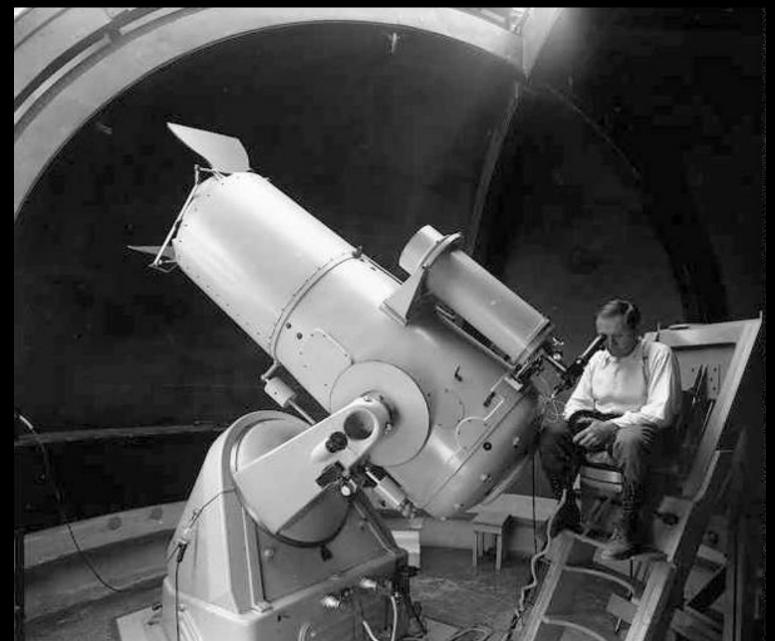


Rare color photograph from 1939 of 18' dome housing g the 18" Schmidt Camera; Inset mid 2006 dome now houses robotic atmospheric turbulence monitoring system. "Caltech Archives" Palomar Observatory's 48-inch Samuel Oschin Telescope **18" Schmidt Camera 1st Light September 5, 1936** 24" mirror 18" correcting plate f/2 fl 914.4 mm (36") 18.2x <u>8 <sup>3</sup>/<sub>4</sub> ° FOV (film-area of 17 full moons )</u>

6 <sup>1</sup>/<sub>4</sub>-inch circles of unexposed film made with the "cookie cutter" in the telescope's darkroom.

Film circle was placed inside a film holder which applied the appropriate spherical curvature so the entire film would be in focus during exposure.

Small access door to place at prime focus. Holder's cap was removed, film was ready for exposure.



### Fritz Zwicky 18" Schmidt Camera 1936 Discovered 120+ Supernovae

Caltech Archives)



200" Palomar Observatory on the Right Caltech Archives-1936



18" focusing on200" Palomar Observatory Caltech Archives-1936



18" focusing on200" Palomar Observatory Caltech Archives-1936

The 18" Schmidt Telescope was funded from the original Rockefeller grant for the construction of the 200-inch.

It was designed by telescope maker Russell W. Porter and engineered at Caltech by John Anderson, Sinclair Smith, and Albert Brower based on a photographic camera invented in 1930 by optician Bernhard Schmidt.

The optics were figured at the Caltech optical shop, while the tube and mounting were manufactured at the Caltech instrument shop. The telescope was finished within a year (1935-1936).

18-inch Schmidt Camera 1<sup>st</sup> instrument at Palomar Observatory Only operational telescope on site between 1936 and 1949.



# Eugene and Carolyn Shoemaker with the 18-inch Schmidt,

Photographed by Jonathan Blair in 1986. (J.Blair/USGS)

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Log 1991 G & C Shoemaker Tech Pan 4415 hypered 6-10 min exposures

Comet Shoemaker–Levy 9 (D/1993 F2)

The comet was discovered by astronomers Carolyn and Eugene M. Shoemaker, and David Levy in 1993.

Shoemaker–Levy 9 (SL9) had been captured by Jupiter and was orbiting the planet at the time. It was located on the night of March 24 in a photograph taken with the 46 cm (18 in) Schmidt telescope at the Palomar Observatory in California.

It was the first active comet observed orbiting a planet, and had probably been captured by Jupiter around 20 to 30 years earlier.

Broke apart in 1992 and struck Jupiter in 1994

# Comet Shoemaker–Levy 9 (D/1993 F2)



Hubble Space Telescope image 1994

In its long and productive life, this instrument has yielded many discoveries, including a large number of asteroids, nearly 50 comets and hundreds of Nova.

The telescope was decommissioned in the mid-1990s and was removed from its dome to make room for a robotic atmospheric turbulence monitoring system in 2006.

The 18" Schmidt Camera was refurbished and is now one of the main exhibits at Palomar Observatory Visitor Center

The 18" Schmidt Camera's Success paved the way for the

Famous: Palomar Observatory 48" Schmidt Telescope



Thanks to a generous gift from the trust of Dr. Helin and her husband Ronald, Palomar Observatory is pleased to announce its new exhibit *Searching the Sky for Dangerous Neighbors* at the Palomar Observatory Visitor Center. Caltech

# The 48" Schmidt Camera

Urging from astronomers Fritz Zwicky & Walter Baade, the 48inch telescope was needed for its wide field of view

Need Sky Surveys to locate & map targets of interest for 200"

The 48-inch Samuel Oschin Telescope is one of the largest Schmidt cameras ever built

Palomar was one of the first observatories in the world to utilize this new technology, which enabled astronomers to survey the sky.

Given the success of the 18-inch (0.46-meter) Schmidt telescope during the 1930s, resources for building the larger 48-inch Schmidt were committed in 1937.

# The 48" Schmidt Camera



Palomar Observatory's 48-inch Samuel Oschin Telescope

# The 48" Schmidt Camera



# The 48" Schmidt Camera

1938 Corning began work on the 72" Pyrex disk/blank. **48**" (1.2m) telescope saw **1**<sup>st</sup> **Light** late September 1948 The concave spherical mirror is 72" (1828 mm/1.2 m) Optical Speed of f/2.5 Plate Holder (2.46) approx. 60x Corrector Plate 49<sup>3</sup>/<sub>4</sub>" (1244 mm) OTA is 20' long 5/16" Steel

Fork Mount combined weight 12 tons

### The 48" Schmidt Camera

Shutters are behind the Corrector Plate

Designed for 10" x 10" & 14"x 14" Photographic Plates Formulated by George Eastman, Eastman Kodak Company

> Optical Speed of f/2.5 (2.46) Field of View (FOV) 36 degrees 14" x 14".

> Manually Guided by two 10-inch refractors

1<sup>st</sup> Photographic Plate November 11,1949

Plates were in wide use by the professional astronomical community as late as the 1990s.

# The 48" Schmidt Camera



Palomar Observatory's 48-inch Telescope one of the most productive survey telescopes ever built rededicated in 1987 to Samuel Oschin Telescope Caltech

# The 48" Schmidt Camera



Astronomer George Abell (1927-1983), a Caltech alumnus and former UCLA professor, at the 48-inch telescope at Palomar... Credit: Caltech Archives

# Palomar Observatory Sky Surveys The most Influential Sky Surveys of their time

1950's Palomar Sky Survey (**POSS-1**) Palomar Observatory & National Geographic

8 years to complete November 11, 1949 →June 20,1956 (99%) Final 1 % completed Dec. 10, 1958

Deepest Photographic record of Northern Sky

1,875 14" x 14" Photographic Plates 937 Pairs

## The 48" Schmidt Camera

1950's Palomar Sky Survey (**POSS-1**) Palomar Observatory & National Geographic

Each region imaged twice

Red Sensitive Kodak 103a-E Plate

Blue Sensitive Kodak 103a-O Plate

Allowing the Color of Celestial objects to be recorded

Down to Magnitude 22

## The 48" Schmidt Camera

1950's Palomar Sky Survey (**POSS-II**) Palomar Observatory & National Geographic

Palomar Sky Survey II 1980's – 1990's

**Better & Faster Film Plates** 

**Upgraded Optics Achromatic Corrector Plate 1979** 

Auto-Guider

3 images of each Region

Blue (IIIaJ. 480 nm), Red (IIIaF, 650 nm), and Near-Infrared (IVN, 850 nm) plates

# The 48" Schmidt Camera

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Palomar Observatory/California Institute of Technology

Actual Log P48 November 1949

# The 48" Schmidt Camera



May 23, 2018 By Sam Pitts

Palomar Observatory's Film Hypering Lab

# The 48" Schmidt Camera

1950's Palomar Sky Survey (**POSS-II**) Palomar Observatory & National Geographic

(POSS II): 3 colors Digitized Sky Survey, 1 Billion stars 50 million galaxies, largest catalog ever

Jean Mueller was hired as 48" Schmidt Camera Night Assistant (July 1985) was the operator for the duration of POSS II and took over 5500 photographic plates, final plate removed from 48" by Jean 6/3/2000

> 1985 48" Schmidt Camera-renamed "Samuel Oschin Telescope"

# The 48" Schmidt Camera



Jean Mueller with Palomar Observatories 48" Schmidt Camera

# The 48" Schmidt Camera



### 48" Observatory Work Room 5/13/2023 Sam Pitts

## The 48" Schmidt Camera



### 48" Observatory Plate Archives Room 5/13/2023 Sam Pitts

#### The 48" Schmidt Camera



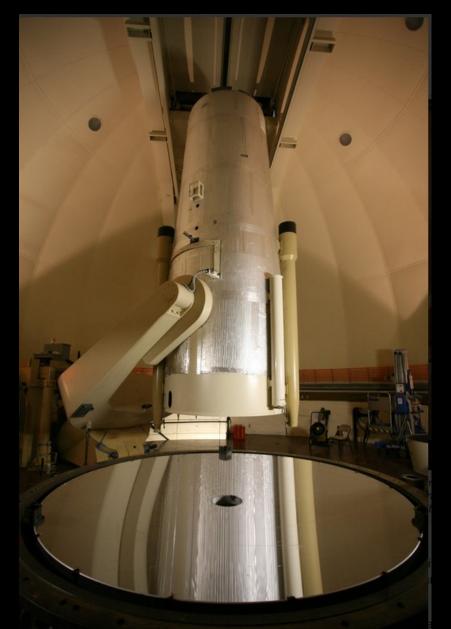
Photo of 14" plates at Palomar Observatory OC 5/13/2023 Sam Pitts

#### The 48" Schmidt Camera



Oschin Telescope as a part of the Second Palomar Observatory Sky Survey (POSS II). (Palomar/Caltech/ In the hild-1980s, the 48-inch Schmidt was upgraded with a new DSS)

#### The 48" Schmidt Camera



Samuel Oschin Telescope 72" Mirror removed for aluminization.

#### The 48" Schmidt Camera

### Huge Leap in Astronomy



### June 3, 2000 The Samuel Oschin Telescope's Last Film Plate

#### The 48" Schmidt Camera

2000 – 2001 Converted to CCD Improved Corrector Plate Wider Range of Wavelengths Match Sensitivity of CCD chips **Auto-Guiders** Automate the Entire Telescope Data collected on:

Performance Wireless Research & Education Network (HPWREN)

#### The 48" Schmidt Camera

1<sup>st</sup> CCD installed (2001) Near Earth Asteroid Tracking **NEAT** CCD Camera 3 - 4k x 4k sensors North to South line Total FOV 3.75 square degrees Upgraded – Automatic Sky Survey every clear Night

300 near-Earth Asteroids and 13 Comets

#### The 48" Schmidt Camera

### 2003 to 2007 Quest 2 Camera

Quasar Equatorial Survey Team Camera

#### 112 CCD's 2400 x 600 pixels each

Arranged in 4 columns of 28 ccd chips with gaps 9.6 square Degrees

Largest Mosaic of CCD chips used in an Astronomical Camera

#### **Palomar Transient Factory**

## Palomar Transient Factory -2008

Mosaic Camera know as CFH12K96

12,288 x 8,192 Pixel Mosaic 12 Chips each 2048 x 4096 15µm pixels

100,000 pixels 200 MB per image

7.8° FOV every 90 seconds

2009-2012

#### **Palomar Transient Factory**

2009 – 2012 The Palomar Transient Factory (PTF):

Fully-automated, Wide-Field Survey

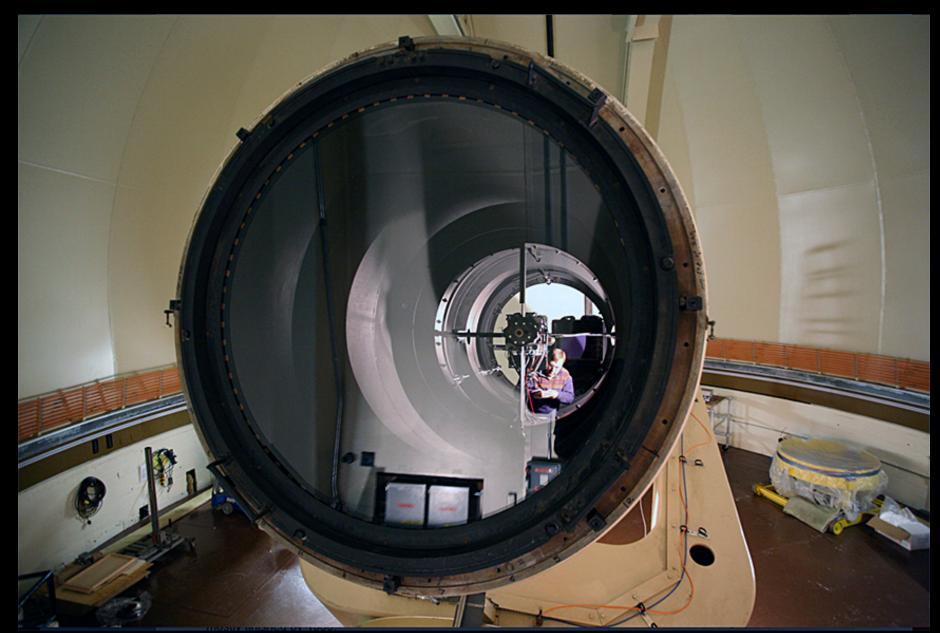
Search for optical transient and variable sources.

PTF discovered 10,000+ of new transient sources

Thousands of new supernovae, novae and cataclysmic variable stars, flaring young stars

1<sup>st</sup> detection of a planet orbiting a young star in Orion.

#### Palomar Transient Factory (PTF)



Installation of the PTF camera - prime focus of the Samuel Oschin Telescope December 2008.

### Intermediate Palomar Transient Factory 2012 – 2017

Mosaic of 11 CCD chips 2048 x 4096 pixels each 7.26° FOV Single exposure 60 seconds 20.5-21.0 Magnitude Based on 2 arcsecond seeing Readout & slew overhead of 40 sec.

Total time per field 100 seconds

Fields repeated with different filters

300 to 400 fields per Night

176 MB per image 50-70 GB per Night

#### **Zwicky Transient Facility**

# Zwicky Transient Facility 1<sup>st</sup> Light October 2017

Zwicky Transient Facility PTF used only 7.3° 48" capable of 47° Time for a New Camera Caltech Developed a New System State of the Art Survey System Camera will have Worlds Widest FOV 6 times the PTF FOV Each Image 235x the area of full moon

Zwicky Transient Facility Survey Strategy

3 High level Programs

1 - Public Surveys 40%

**2-** ZTF Collaboration Surveys 40%

**3-** Caltech Surveys 20%

Each Program has multiple Sub-Programs

Zwicky Transient Facility 2017 – Present

### Fall 2018

ZTF Observing System was fully operational Generating significant numbers of transient alerts in real-time for the US astronomical community every clear Observing Night

Formal survey operations began on March 20, 2018

#### **Zwicky Transient Facility**

ZTF took this "first-light" image on Nov. 1, 2017 47 Square Degrees

#### **Zwicky Transient Facility**



Working f ratio f/2.46 (f/2.5) 48" = 1219.2mm fl-2999.232 approximately 60x CCD 6144 x 6160 15µm full well 350,000 e- FOV 47°

#### **Zwicky Transient Facility**

Publications of the Astronomical Society of the Pacific, 132:038001 (26pp), 2020 March

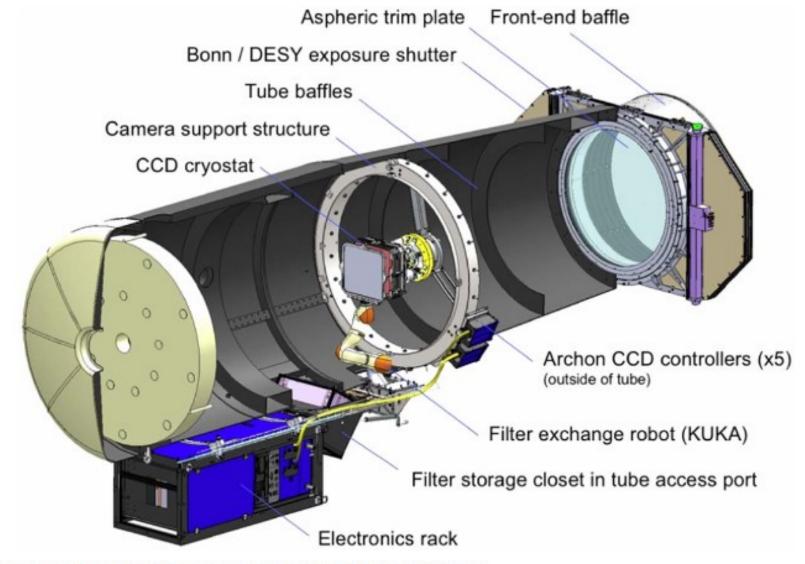


Figure 1. Cutaway view of the Samuel Oschin Telescope highlighting new ZTF subsystems. (A color version of this figure is available in the online journal.)

#### **Zwicky Transient Facility**

#### Publications of the Astronomical Society of the Pacific, 132:038001 (26pp), 2020 March

Dekany et al.

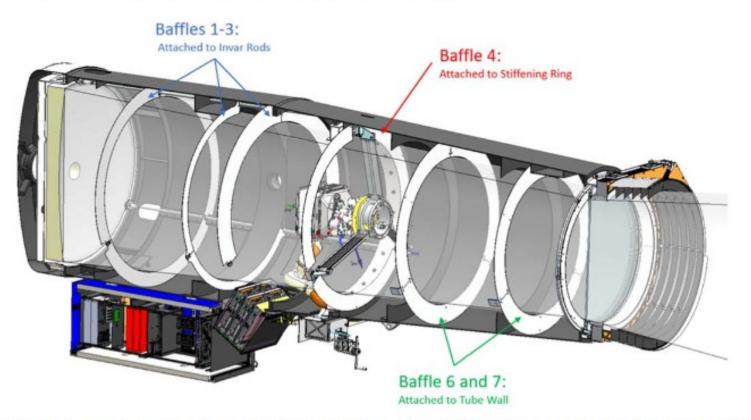
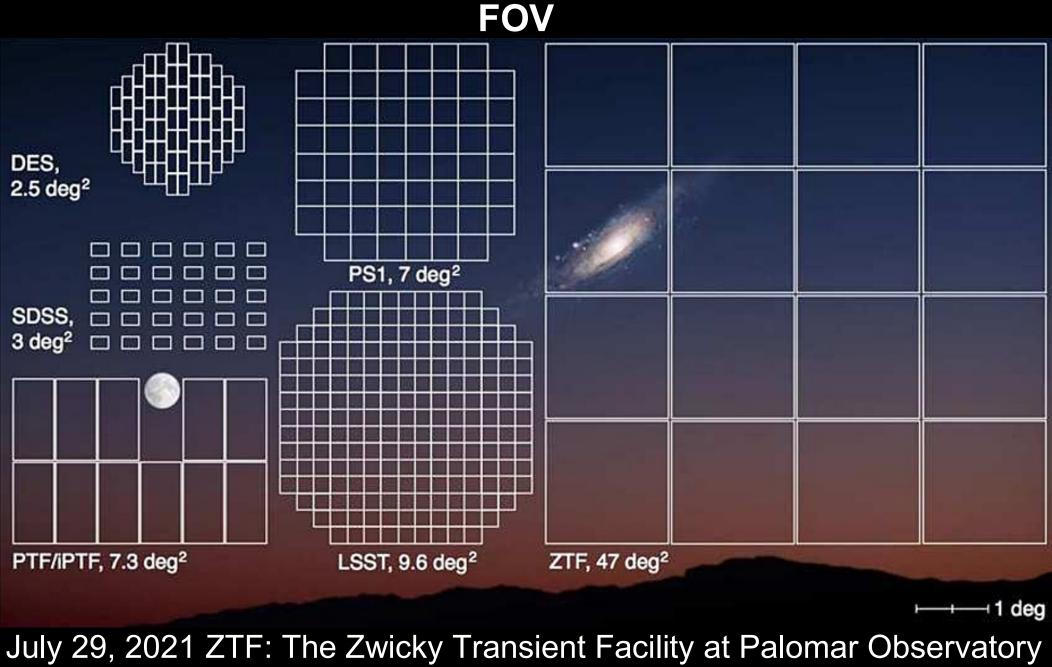


Figure 10. ZTF telescope tube baffling layout consists of 7 concentric baffles arranged so that off-axis scattered (e.g., moon) light cannot reach the primary mirror without scattering at least twice from blackened surfaces.

#### **Zwicky Transient Facility**



#### **Zwicky Transient Facility**

#### ZTF will survey an order of magnitude faster than PTF.

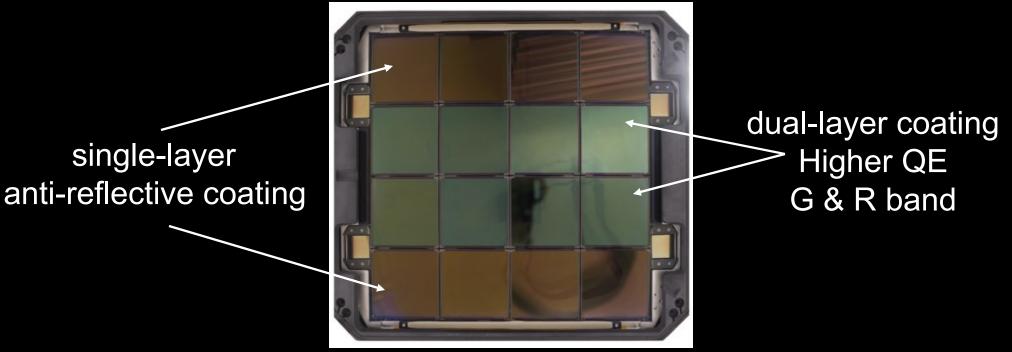
	PTF	ZTF
Active Area	7.26 deg <sup>2</sup>	47 deg <sup>2</sup>
Overhead Time	46 sec	<15 sec
Optimal Exposure Time	60 sec	30 sec
Relative Areal Survey Rate	1x	15.0x
Relative Volumetric Survey Rate	1x	12.3x

MOSAIC 12k

## 3750 deg<sup>2</sup>/hour $\Rightarrow$ $3\pi$ survey in 8 hours >250 observations/field/year for uniform survey New ZTF camera: 16 6k x 6k e2v CCDs Existing PTF camera

#### **Zwicky Transient Facility**

## Filter transmission for the ZTF *g*, *r*, and *i*-band filters (blue, orange, and red lines).



Palomar Observatory/California Institute of Technology

Four  $2k \times 2k$  CCD's are located on the perimeter of the mosaic; one serves as a guider while the remaining three control tip, tilt, and focus. North is up and east is left.

#### **Zwicky Transient Facility**

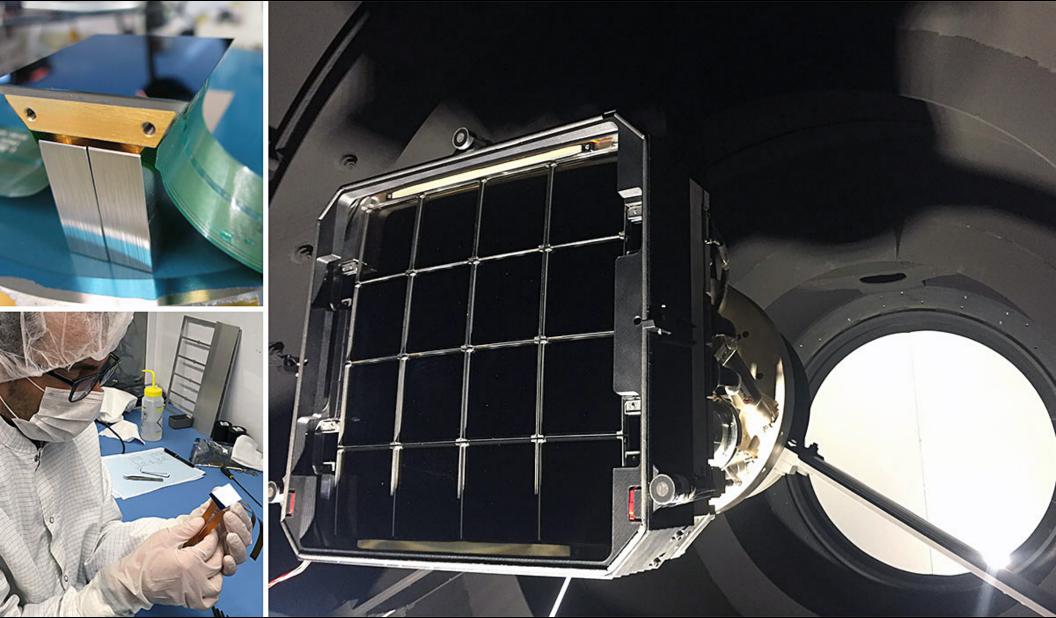
Field of view	47 sq. degrees
Detectors	16 e2v 6k x 6k CCD231-C6
Pixel size	15 microns
Well Depth	350,000 -е
Pixel scale	1.0"/pixel
Median Delivered Image Quality	2.0" FWHM
Exposure time	30 sec
Readout time	10 sec
Time Between Exposures	15 sec
Median 30 sec ( R band )	20.4 mag. (all lunar phases)
Filters	ZTF g, ZTF r, ZTF i
Areal survey rate	3750 square degrees/hour
CCD Cooling	l60K (-171.67° F)

#### **Zwicky Transient Facility**



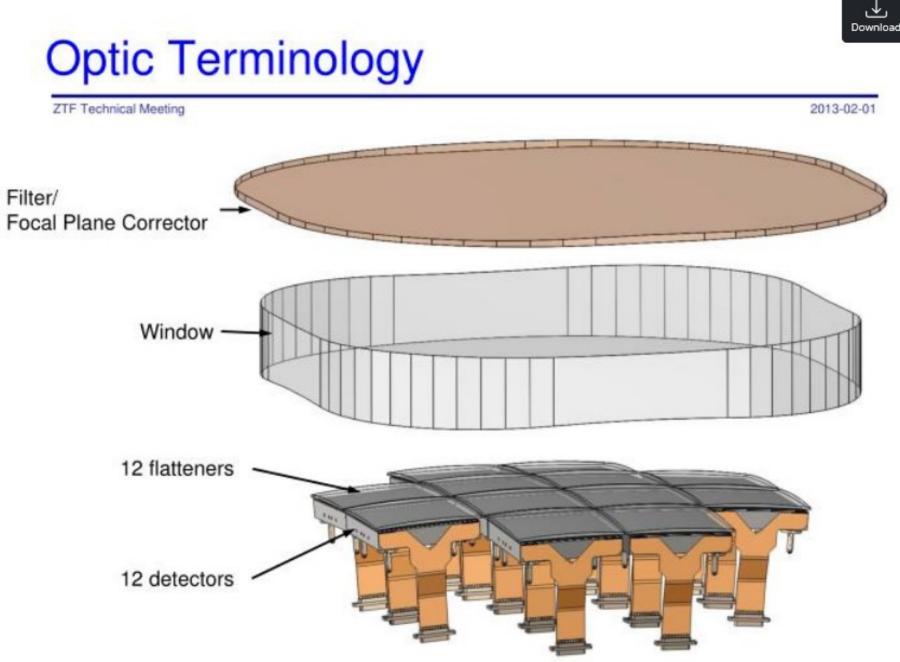
#### New ZTF Camera 16 6,000 x 6,000 e2v- CCD231-C6

#### **Zwicky Transient Facility**



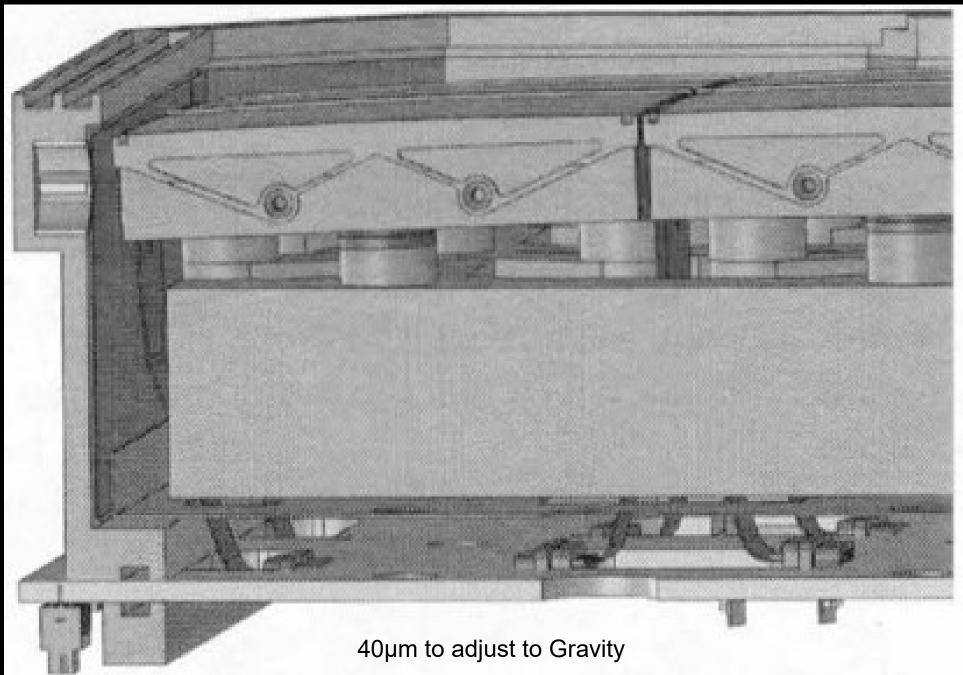
#### New ZTF 47° FOV Camera in 48" OTA at Focus Plane

### **Zwicky Transient Facility**



Palomar Observatory/California Institute of Technology

#### **Zwicky Transient Facility**



#### **Zwicky Transient Facility**



#### **ZTF** Filter

#### **Zwicky Transient Facility**

#### 16 x 16 CCD and Robotic Arm for Filters

#### Zwicky Transient Facility CCD Data

San Diego Supercomputing Center (SDSC)

Data to the IPAC servers in Pasadena via HPWREN High Performance Wireless Research & Education Network

Caltech's Infrared Processing and Analysis Center (IPAC),

The ZTF Science Data System (ZSDS) is housed at IPAC

Converted to 32 bit floating point FITS & Calibrated

 $650 \rightarrow 900$  exposures per night

Archive Data  $\rightarrow$  4 terabytes each night

#### **Zwicky Transient Facility**

#### Programs

ZTF is conducting the two Public surveys

# Northern Sky Survey and a Galactic Plane Survey.

Mid-scale Innovations Program (MSIP):

Galactic Plane Survey is a nightly survey

Northern Sky Survey is a three-day cadence survey Records Comets & Centaurs

(bodies between Jupiter & Neptune)

Each field is visited twice  $\rightarrow$  g-band & r-band

#### **Zwicky Transient Facility**

1<sup>st</sup> 5 Years 2018-2023

ZTF discovered and classified over 8000 supernovae

Over 3000 Type 1a supernovae (binaries) events

Hundreds of near-Earth asteroids

Ten's of rare transients like Tidal Disruption events (when a star is violently ripped apart by the gravity of a black hole.)

15,000+ observations of Comets

### Zwicky Transient Facility Currently 2023-2025

Caltech continues to press the Edge of the Envelope Improvements & Upgrades continue to emerge

Dome, Telescope and Electronic Upgrades Allow the 48" Samuel Oschin Telescope to slew and settle between adjacent fields separated by 7° during the CCD readout time of 10 seconds

Rumored readout times of 8.66 seconds can be achieved.

Fastest Observing Cadence 38.98 ± 0.14 seconds

Autoguider: ZWO ASI1600MM CMOS sensor

#### **Zwicky Transient Facility**

## If it could ignore Daylight ZTF could repeat the entire POSS survey in one day.